LONG RANGE CORRELATIONS;EVENT SIMULATION AND PARTON PERCOLATION

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BNL Glasma Workshop, May 10-12

Introduction to long range correlations
STAR data and string like models
Clustering of color sources
Scales of pp and Pb Pb collisions
Similarities between CGC and percolation
Results on b for pp and AA
Conclusions

LONG RANGE CORRELATIONS

• A measurement of such correlations is the backward–forward dispersion

$$D^2_{FB} = < n_B n_F > - < n_B > < n_F >$$

where $n_B(n_F)$ is the number of particles in a backward (forward) rapidity

$$D_{FB}^{2} = < N > (< n_{1B}n_{1F} > - < n_{1B} > < n_{1F} >) + (< N^{2} > - < N >^{2}) < n_{1F} > < n_{1B} >$$

$$- | B | | | | | | | | | | | | |$$

$$\wedge m$$

<N> number of collisions: <n_{1B}>,<n_{1F}> F and B multiplications in one collision

• In a superposition of independent sources model, D_{BF}^2 is proportional to the fluctuations (D_N^2) on the number of independent sources (It is assumed that Forward and backward are defined in such a way that there is a rapidity window $\Delta_n \ge 1.0$ to eliminate short range correlations).

Correlation parameter $< n_B >= a + b n_F$ with $b \equiv D_{BF}^2 \, / \, D_{FF}^2$

• b in pp increases with energy. In hA increases with A also in AA,increases with centrality

The dependence of b with rapidity gap is quite interesting, remaining flat for large values of the rapidity window.

Existence of long rapidity correlations at high density

Correlation Parameter b

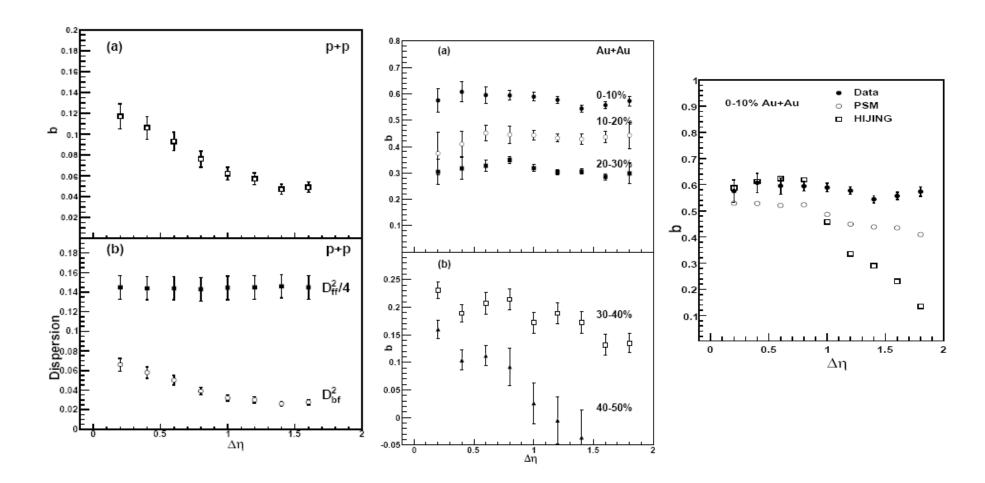
I Situation: Symmetrica

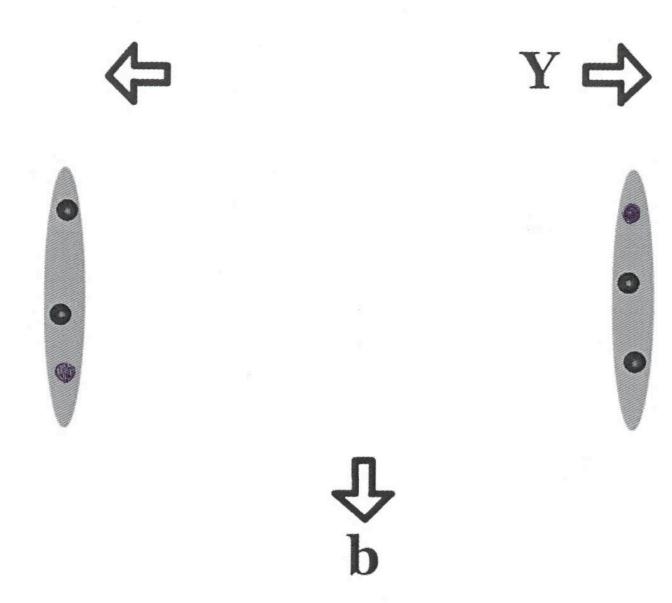
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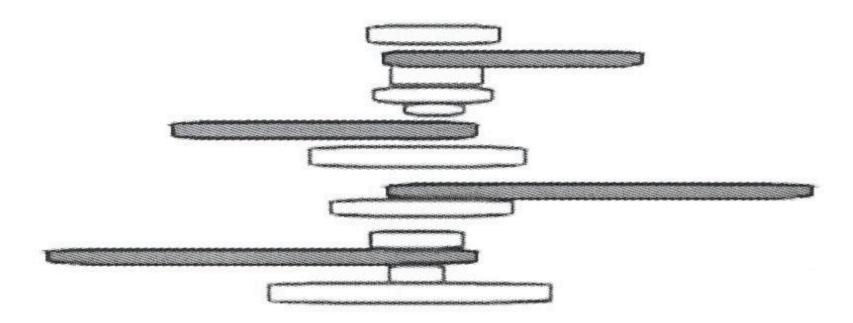
$$b \equiv \frac{1}{1 + \frac{K}{\langle n_F \rangle}}$$

• 1/K is the squared normalized fluctuations on effective number of strings(clusters)contributing to both forward and backward intervals

The heigth of the ridge structure is proportional to n/k







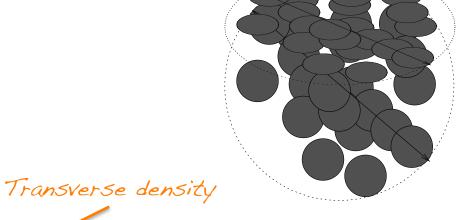
--The strings must be extended in both hemispheres, otherwise either they do not obtained LRC(Heijing)or they have to include parton interactions(PACIAE).PACIAE reproduces well b for central but not for peripheral --Without parton interactions the length of the LRC is the same in pp than AA(modified wounded model of Bzdak)

CLUSTERING OF COLOR SOURCES

- Color strings are stretched between the projectile and target
- Strings = Particle sources: particles are created via sea qqbar production in the field of the string
- Color strings = Small areas in the transverse space filled with color field created by the colliding partons
- With growing energy and/or atomic number of colliding particles, the number of sources grows
- So the elementary color sources start to overlap, forming clusters, very much like disk in the 2-dimensional percolation theory
- In particular, at a certain critical density, a macroscopic cluster appears, which marks the percolation phase transition

(N. Armesto et al., PRL77 (96); J.Dias de Deus et al., PLB491 (00); M. Nardi and H. Satz(98).

• How?: Strings fuse forming clusters. At a certain critical density η_c (central PbPb at SPS, central AgAg at RHIC, central pp at LHC) a macroscopic cluster appears which marks the percolation phase transition (second order, non thermal).



 $\eta = N_{st} \frac{S_1}{S_A} \;, \quad S_1 = \pi r_0^2, \quad r_0 = 0.2 \quad \text{fm}, \quad \eta_c = 1.1 \div 1.2.$

• Hypothesis: clusters of overlapping strings are the sources of particle production, and central multiplicities and transverse momentum distributions are little affected by rescattering.

Particle density from string cluster $\mu_n = \sqrt{\frac{nS_n}{S_1}}\mu_1 \; ; \; < p_T^2 >_n = \sqrt{\frac{nS_1}{S_n}} < p_T^2 >_1$

Energy-momentum of the cluster is the sum of the energy-momentum of each string. As the individual color field of the individual string may be oriented in an arbitrary manner respective to one another, $Q_n^2 = nQ_1^2$

At high densities

•
$$<\mu>_n=nF(\eta)<\mu>_1< p_T^2>_n=\frac{< p_T^2>_1}{F(\eta)}$$

•
$$F(\eta) = \sqrt{\frac{1-e^{-\eta}}{\eta}}$$
, $\eta = N_S \frac{\pi r_0^2}{S_A}$

• r_0 is the transverse size of a single string $\simeq 0.2$ fm.

Scales of pp and AA

Why Protons?

In String Percolation...

$$\eta_{AA} = \left(\frac{r}{R}\right)^2 \overline{N}^s \cong \frac{N_A^{4/3}}{N_A^{2/3}} \left(\frac{r}{R_p}\right)^2 \overline{N}_p^s$$

$$\eta_{AA}(s) = N_A^{2/3} \eta_{pp}(s) \quad \text{and} \quad \overline{N} \sim s^{2/7}$$

$$\eta_{PbPb}(\sqrt{s}) \cong 20 GeV - 200 GeV$$
 $\eta_c \approx 1.15$
 $\eta_{PP}(\sqrt{s}) \cong 6 TeV - 14 TeV$
LHC

Comments

Energy momentum conservation leads to increase in rapidity (length) of string

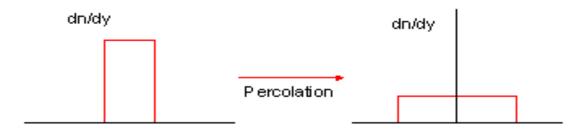
- Data (RHIC)
- FB Correlations YES: SO LONG?

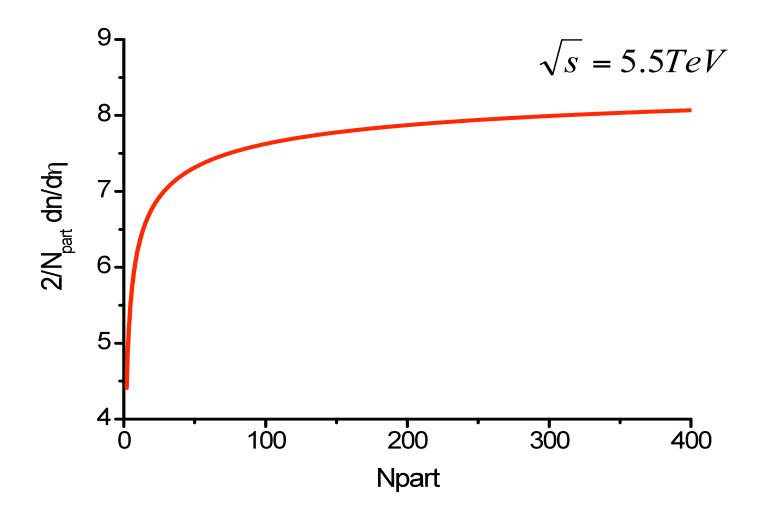
Colour Flux Tube: OK

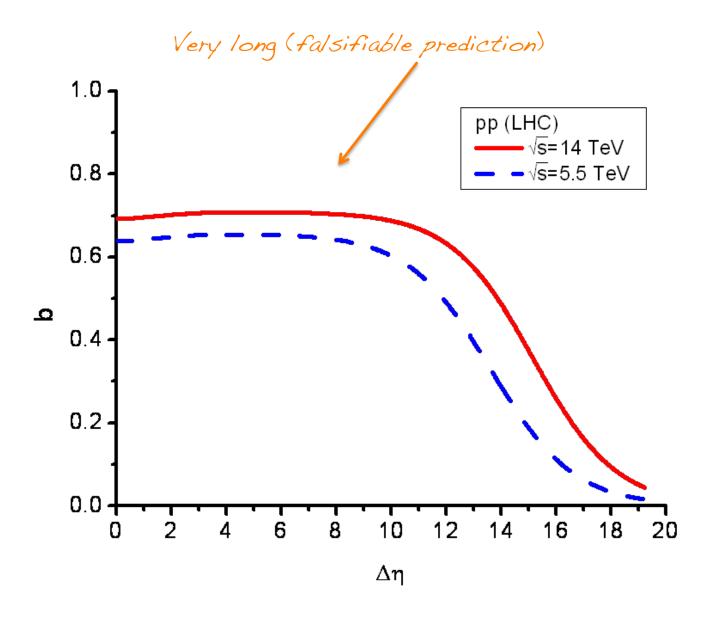
Strings : TOO SHORT

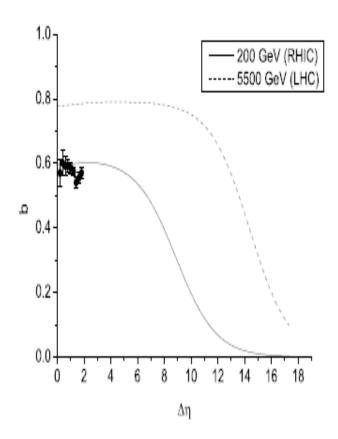
One String: $x^+ = x^- = x = 1/\sqrt{s} \Rightarrow \Delta y_1$

 N^s Strings: $\Delta y_{N_s} = \Delta y_1 + 2 \ln(N_s)$









 $r_0^2 F(\eta)$ Qualitative dictionary PERC-CGC

Transverse size

CGC

$$r_0^2 F(\eta)$$
 $\frac{1}{Q}$

Effective number of clusters

$$< N > = \frac{(1 - exp(-\eta))R_A^2}{F(\eta) r_0^2} = (1 - exp(\eta))^{1/2} \sqrt{\eta} \left(\frac{R_A}{r_0}\right)^2$$

low density

$$\eta \left(rac{R_A}{r_0}
ight)^2$$
 , $N_A^{4/3}$, $exp(2\lambda\gamma)$

high density

$$\sqrt{\eta} \left(\frac{R_A}{r_o}\right)^2, N_A, \exp(\lambda \gamma)$$

$$CGC \; rac{1}{lpha_S} R_A^2 Q_S^2 \; , \, N_A \; , \, exp(\lambda \gamma \;)$$

rapidity extension

$$\Delta y_N = \Delta y_I + 2lnN_S \quad ,$$

$$lnN_A$$
, lns

$$CGC \quad \frac{1}{\alpha_s}$$

$$ln N_A$$
 , $ln s$

1/k = normalized fluctuation of eff number of strings

$$K = \frac{\langle N \rangle^2}{\langle N^2 \rangle - \langle N \rangle^2}$$

low density

$$k \rightarrow \infty$$

high density

$$k \rightarrow \infty$$

$$k = \frac{\langle N \rangle}{(1 - exp(\eta))^{\frac{3}{2}}}$$

high density

$$\sqrt{\eta} \left(\frac{R_A}{r_0} \right)^2$$
 , N_A , $exp(\lambda \gamma)$

low density

$$\frac{1}{\sqrt{\eta}} \left(\frac{R_A}{r_0} \right)^2$$

$$CGC \quad k=R_S^2Q_S^2$$
, N_A , $exp(\lambda \gamma)$

MULTIPLICITY DISTRIBUTIONS

NEGATIVE BINOMIAL

low density
$$k \to \infty$$
 , $k_0 \to \infty$ Poisson

high density
$$k \to \infty$$
 , $k_0 \to 1$ Bose-Einstein

$$CGC$$
 $k_0 = 1$ $B:E$, $k=\langle N \rangle$

k first decreases with density (energy)

Above an energy(density) k increases

⇒ Multiplicity distributions (normalized,

i.e. $< n > P_n$ as a function of n / < n > will be narrower (Quantum Optical prediction)

$$b = \frac{1}{1 + \frac{d}{(1 - e^{-\eta})^{\frac{3}{2}}}}$$

low density

$$b \rightarrow 0$$

high density (energy)

$$b \rightarrow \frac{1}{1+d}$$

CGC

$$b = \frac{1}{1 + \alpha_S^2 c}$$

high density (energy)

$$b \rightarrow 1$$

Conclusions

- ---For pp at LHC are predicted the same phenomena observed at RHIC in Au-Au
- ---Normalized multiplicity distributions will be narrower
- ---Long range correlations extended more than 10 units of rapidity at LHC.Large LRC in pp extended several units of rapidity.
- ---Large similarities between CGC and percolation of strings.Similar predictions corresponding to similar physical picture.Percolation explains the transition low density-high density

a few notes

- · Close relation between Glasma and SPM
 - (however) no magnetic fields here
- · No fundamental theoretical basis
 - But, clear physical picture and straightforward calculational framework
 - Good testbed for qualitative features